

A. Ivashkevich, A. Buryy, E. Ovsiyuk, V. Balan, V. Kisel, V. Red'kov
 B. I. Stepanov Institute of Physics,
 National Academy of Sciences of Belarus, Minsk, Belarus

ON THE MATRIX EQUATION FOR A SPIN 2 PARTICLE IN PSEUDO-RIEMANNIAN SPACE-TIME

After the study by Pauli and Fierz [1, 2], the theory of massive and massless fields with spin 2 has always attracted much attention [3–13]. Most of the studies were performed in the framework of 2-nd order differential equations. It is known that many specific difficulties may be avoided if from the very beginning we start with 1-st order systems. Apparently, the first systematic study of the theory of spin 2 fields within the first order formalism was done by F. I. Fedorov [4]. It turns out that this description requires a field function with 3 independent components. This theory was re-discovered and improved by Regee [5]. In the present paper we develop the theory of the spin 2 field, in both massive and massless variants, starting from the matrix equation in Minkowski space-time and extending it to the generally covariant theory within the Tetrad-Weyl-Fock-Ivanenko tetrad method.

We start with the known system of the first order equations for a massive spin 2 particle:

$$\begin{aligned} \partial^a \Phi_a &= m\Phi, \quad \frac{1}{2} \partial_a \Phi - \frac{1}{3} \partial^b \Phi_{(ab)} = m\Phi_a, \\ \frac{1}{2} \left(\partial^k \Phi_{[ka]b} + \partial^k \Phi_{[kb]a} - \frac{1}{2} g_{ab} \partial^k \Phi_{[kn]} \right) + \left(\partial_a \Phi_b + \partial_b \Phi_a - \frac{1}{2} g_{ab} \partial^k \Phi_k \right) &= m\Phi_{(ab)}, \\ \partial_a \Phi_{(bc)} - \partial_b \Phi_{(ac)} + \frac{1}{3} \left(g_{bc} \partial^k \Phi_{(ak)} - g_{ac} \partial^k \Phi_{(bk)} \right) &= m\Phi_{[ab]c}, \end{aligned} \quad (1)$$

where the field variables are scalar, vector, symmetric 2-rank tensor, and 3-rank skew-symmetric in two first indices tensor, $m = iM$. By excluding the vector and the 3-rank tensor, we obtain the 2-nd order equations with respect to the scalar and symmetric tensor:

$$\Phi = 0, \quad (\square + M^2)\Phi_{(ab)} = 0, \quad \Phi_{(ab)} = \Phi_{(ba)}, \quad \Phi^a_a = 0, \quad \partial^k \Phi_{(ka)} = 0. \quad (2)$$

In massless case, the first order system reads

$$\begin{aligned}
\partial^a \Phi_a &= 0, \quad \frac{1}{2} \partial_a \Phi - \frac{1}{3} \partial^b \Phi_{(ab)} = \Phi_a, \\
\frac{1}{2} \left(\partial^k \Phi_{[ka]b} + \partial^k \Phi_{[kb]a} - \frac{1}{2} g_{ab} \partial^k \Phi_{[kn]} \right) + \left(\partial_a \Phi_b + \partial_b \Phi_a - \frac{1}{2} g_{ab} \partial^k \Phi_k \right) &= 0, \\
\partial_a \Phi_{(bc)} - \partial_b \Phi_{(ac)} + \frac{1}{3} \left(g_{bc} \partial^k \Phi_{(ak)} - g_{ac} \partial^k \Phi_{(bk)} \right) &= \Phi_{[ab]c}. \quad (3)
\end{aligned}$$

From (3) we derive the 2-nd order equations for the massless field:

$$\begin{aligned}
\frac{1}{2} \square \Phi - \frac{1}{3} \partial^k \partial^l \Phi_{(kl)} &= 0, \\
(\partial_a \partial_b + \frac{1}{2} g_{ab} \square) \Phi - \frac{1}{4} g_{ab} \square \Phi_c + \square \Phi_{(ab)} - \partial_a \partial^l \Phi_{(bl)} - \partial_b \partial^l \Phi_{(al)} &= 0. \quad (4)
\end{aligned}$$

Massless equations have a class of gauge solutions:

$$\bar{\Phi} = \partial^l L_l, \quad \bar{\Phi}_{(ab)} = \partial_a L_b + \partial_b L_a - \frac{1}{2} g_{ab} \partial^l L_l, \quad (5)$$

where $L_l(x)$ stands for an arbitrary 4-vector. These special states do not contribute to physically observable quantities, like the energy-momentum tensor. The concomitant gauge components are as follows:

$$\bar{\Phi}_a = \frac{1}{3} \partial_a \partial^l L_l - \frac{1}{3} L_a, \quad \bar{\Phi}_{[ab]c} = \partial_c (\partial_a L_b - \partial_b L_a) - \frac{g_{cb} \partial_a - g_{ca} \partial_b}{3} \partial^l L_l + \frac{g_{cb} L_a - g_{ca} L_b}{3}. \quad (6)$$

The system (1) can be re-written in equivalent block form

$$\begin{aligned}
\partial_a (G^a)_{(0)}^k \Phi_k &= m \Phi_{(0)}, \quad \partial_a \left\{ \frac{1}{2} (\Delta^a)_k^{(0)} \Phi_{(0)} - \frac{1}{3} (K^a)_k^{(mn)} \Phi_{mn} \right\} = m \Phi_k, \\
\partial_a \left\{ \frac{1}{2} (B^a)_{(cd)}^{[mn]l} \Phi_{mnl} + (\Lambda^a)_{(dc)}^k \Phi_k \right\} &= m \Phi_{dc}, \quad \partial_a \left\{ (F^a)_{[kb]c}^{(mn)} \Phi_{mn} \right\} = m \Phi_{kbc}, \quad (7)
\end{aligned}$$

The corresponding matrix equation

$$(\Gamma^a \frac{\partial}{\partial x^a} - m) \Psi(x) = 0, \quad \Psi = \{H; H_1; H_2; H_3\} \quad (8)$$

is extended to the Riemannian space-time in accordance with the tetrad method. In a space-time with given metric, we fix a tetrad:

$$dS^2 = g_{\alpha\beta}(x)dx^\alpha dx^\beta, \quad g_{\alpha\beta}(x) \rightarrow e_{(a)\alpha}(x), \quad g_{\alpha\beta}(x) = \eta^{ab} e_{(a)\alpha}(x)e_{(b)\beta}(x), \quad (9)$$

and then the generalized form gets written as follows

$$\left[\Gamma^\alpha(x) \left(\frac{\partial}{\partial x^\alpha} + \Sigma_\alpha(x) \right) - m \right] \Psi(x) = 0, \quad (10)$$

where the local matrices $\Gamma^\alpha(x)$ are determined with the use of the tetrad

$$\Gamma^\alpha(x) = e_{(a)}^\alpha(x) \Gamma^a = \begin{pmatrix} 0 & G^\alpha(x) & 0 & 0 \\ \frac{1}{2} \Delta^\alpha(x) & 0 & -\frac{1}{3} K^\alpha(x) & 0 \\ 0 & \Lambda^\alpha(x) & 0 & \frac{1}{2} B^\alpha(x) \\ 0 & 0 & F^\alpha(x) & 0 \end{pmatrix}, \quad (11)$$

and connection $\Sigma_\alpha(x)$ is defined by relations

$$\Sigma_\alpha(x) = J^{ab} e_{(a)}^\beta(x) e_{(b)\beta;\alpha}(x) = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & (\Sigma_1)_\alpha & 0 & 0 \\ 0 & 0 & (\Sigma_2)_\alpha & 0 \\ 0 & 0 & 0 & (\Sigma_3)_\alpha \end{pmatrix}, \quad (12)$$

where $\Sigma_i(x) = J_i^{ab} e_{(a)}^\beta(x) e_{(b)\beta;\alpha}(x)$, $i=1,2,3$; and $J_1^{ab}, J_2^{ab}, J_3^{ab}$ stand for the generators for the tensors $\Phi_k, \Phi_{(mn)}, \Phi_{[mn]l}$. The equation (10) can be presented by using the Ricci rotation coefficients

$$\left[\Gamma^c \left(e_{(c)}^\alpha(x) \frac{\partial}{\partial x^\alpha} + \frac{1}{2} J^{ab} \gamma_{abc} \right) - m \right] \Psi(x) = 0. \quad (13)$$

In block form, eq. (13) reads

$$G^\alpha(x)[\partial_\alpha + (\Sigma_1)_\alpha]H_1 = mH, \quad \frac{1}{2}\Delta^\alpha(x)\partial_\alpha H - \frac{1}{3}K^\alpha(x)[\partial_\alpha + (\Sigma_2)_\alpha]H_2 = mH_1,$$

$$\Lambda^\alpha(x)[\partial_\alpha + (\Sigma_1)_\alpha]H_1 + \frac{1}{2}[\partial_\alpha + (\Sigma_3)_\alpha]H_3 = mH_2, \quad F^\alpha(x)[\partial_\alpha + (\Sigma_2)_\alpha]H_2 = mH_3.$$

In the massless case, the system slightly changes:

$$G^\alpha(x)[\partial_\alpha + (\Sigma_1)_\alpha]H_1 = 0, \quad \frac{1}{2}\Delta^\alpha(x)\partial_\alpha H - \frac{1}{3}K^\alpha(x)[\partial_\alpha + (\Sigma_2)_\alpha]H_2 = H_1,$$

$$\Lambda^\alpha(x)[\partial_\alpha + (\Sigma_1)_\alpha]H_1 + \frac{1}{2}[\partial_\alpha + (\Sigma_3)_\alpha]H_3 = 0, \quad F^\alpha(x)[\partial_\alpha + (\Sigma_2)_\alpha]H_2 = H_3,$$

but its physical content is completely different. In particular, let us detail tetrad representation for the gauge solutions:

$$\bar{\Phi} = \nabla_\alpha L^\alpha(x) \Rightarrow \bar{\Phi} = e^{(c)\alpha} \partial_\alpha L_{(c)} + e_{(c);\alpha}^\alpha L^{(c)},$$

$$\bar{\Phi}_{(\alpha\beta)} = \nabla_\alpha L_\beta + \nabla_\beta L_\alpha - \frac{1}{2}g_{\alpha\beta}(x)\nabla_\rho \Lambda^\rho \Rightarrow$$

$$\bar{\Phi}_{(ab)} = -\left(\gamma_{[ca]b} + \gamma_{[cb]a}\right)L^{(c)} + e_{(a)}^\alpha \partial_\alpha \Lambda_{(b)} + e_{(b)}^\alpha \partial_\alpha \Lambda_{(a)} - \frac{1}{2}g_{ab}\bar{\Phi}. \quad (14)$$

The concomitant gauge components are determined by the formulas

$$\bar{H}_1 = \frac{1}{2}\Delta^\alpha(x)\partial_\alpha \bar{H} - \frac{1}{3}K^\alpha(x)[\partial_\alpha + (\Sigma_2)_\alpha]\bar{H}_2, \quad \bar{H}_3 = F^\alpha(x)[\partial_\alpha + (\Sigma_2)_\alpha]\bar{H}_2. \quad (15)$$

The covariant equation is symmetric under the local Lorentz group, in accordance with the following relations

$$\Psi'(x) = S(x)\Psi(x), \quad S(x)\Gamma^\alpha(x)S^{-1}(x) = \Gamma'^\alpha(x),$$

$$S(x)\Sigma_\alpha(x)S^{-1}(x) + S(x)\frac{\partial}{\partial x^\alpha}S^{-1}(x) = \Sigma'_\alpha, \quad (16)$$

where the prime indicates that quantities are determined with the use of the primed tetrad related to initial one by the local Lorentz transformation $e_{(a')}^\sigma(x) = L_a^b(x)e_{(b)}^\sigma(x)$. With respect to the coordinate transformation, the field function Ψ behaves as a scalar, $x^\alpha \rightarrow x'^\alpha$, $\Psi(x) = \Psi'(x')$.

References

1. Fierz, M. On relativistic wave equations for particles of arbitrary spin in an electromagnetic field / M. Fierz, W. Pauli // Proc. Roy. Soc. London. A. – 1939. – Vol. 173. – P. 211–232.
2. Pauli, W. Über relativistische Feldgleichungen von Teilchen mit beliebigem Spin im elektromagnetischen Feld / W. Pauli, M. Fierz // Helv. Phys. Acta. – 1939. – Vol. 12. – P. 297–300.
3. Fradkin, E. S. To the theory of particles with higher spins / E. S. Fradkin // Journal of Experimental and Theoretical Physics. – 1950. – Vol. 20, № 1. – P. 27–38.
4. Fedorov, F. I. To the theory of particles with spin 2 / F. I. Fedorov // Proceedings of Belarus State University. Ser. Phys.-Math. – 1951. – Vol. 12. – P. 156–173.
5. Regge, T. On properties of the particle with spin 2 / T. Regge // Nuovo Cimento. – 1957. – Vol. 5, № 2. – P. 325–326.
6. Krylov, B. V. Equations of the first order for graviton / B. V. Krylov, F. I. Fedorov // Doklady of the National Academy of Sciences of Belarus. – 1967. – Vol. 11, № 8. – P. 681–684.
7. Fedorov, F. I. Equations of the first order for gravitational field / F. I. Fedorov // Doklady of the Academy of Sciences of USSR. – 1968. – Vol. 179, № 4. – P. 802–805.
8. Bogush, A. A. On matrices of the equations for spin 2 particles / A. A. Bogush, B. V. Krylov, F. I. Fedorov // Proceedings of the National Academy of Sciences of Belarus. Physics and Mathematics Series. – 1968. – № 1. – P. 74–81.
9. Fedorov, F. I. The first order equations for gravitational field in vacuum / F. I. Fedorov, A. A. Kirilov // Acta Physica Polonica. B. 1976. – Vol. 7, № 3. – P. 161–167.
10. Kisel, V. V. On relativistic wave equations for a spin 2 particle / V. V. Kisel // Proceedings of the National Academy of Sciences of Belarus. Physics and Mathematics Series. – 1986. – № 5. – P. 94–99.
11. On equations for a spin 2 particle in external electromagnetic and gravitational fields / A. A. Bogush [et al.] // Proceedings of the National Academy of Sciences of Belarus. Physics and Mathematics Series. – 2003. – № 1. – P. 62–67.
12. Red'kov, V. M. Graviton in a curved space-time background and gauge symmetry / V. M. Red'kov, N. G. Tokarevskaya, V. V. Kisel // Non-linear Phenomena in Complex Systems. – 2003. Vol. 6, № 3. – P. 772–778.

13. Contribution of gauge degrees of freedom in the energy-momentum tensor of the massless spin 2 field / V. V. Kisel [et al.] // Proceedings of the National Academy of Sciences of Belarus. Physics and Mathematics Series. – 2015. – № 2. – P. 58–63.